

La maîtrise des vibrations dans les constructions civiles
et les structures de précision

André Preumont
Laboratoire des Structures Actives
Université Libre de Bruxelles

AVE 2016
Blois

ULB

1

Tacoma bridge flutter instability (1940)
[Wind speed: 40 mph]

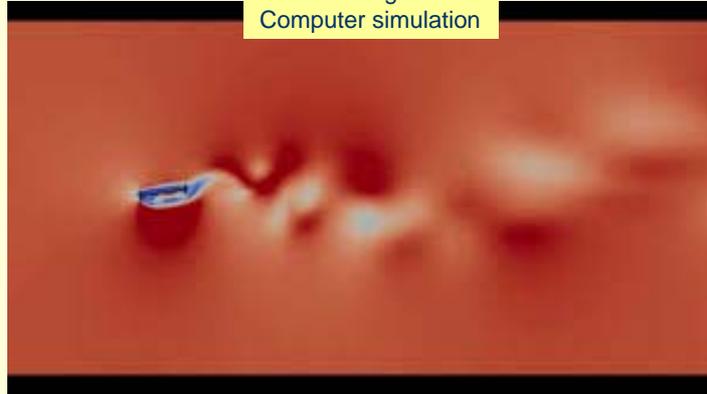


Instability resulting from the interaction between the deck vibrations
and the unsteady aerodynamic forces acting on the deck

ULB

2

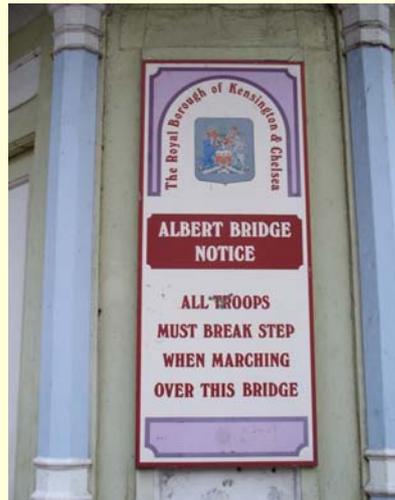
Tacoma bridge Flutter
Computer simulation



ULB

3

Pedestrian-induced vibrations of bridges



4

ULB

Pedestrian-induced vibrations of footbridges



Human Activity	Frequency Range (Hz)
Walking	1,6-2,4
Running	2,0-3,5
Jumping	1,8-3,4

Seriata footbridge:
 8 walking pedestrians: 1.8 m/s²
 4 joggers: 4 m/s²

5

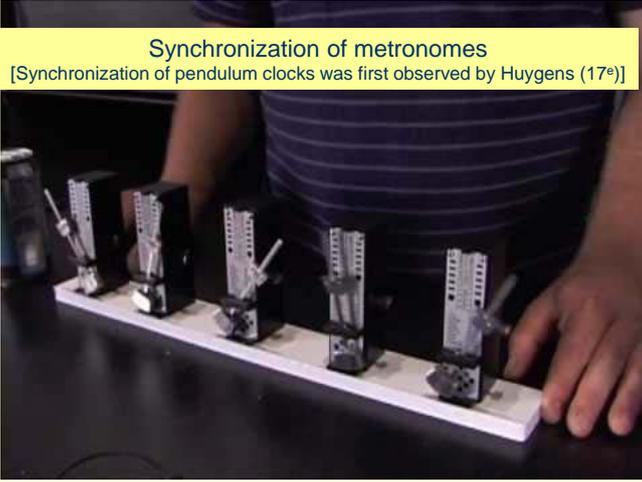
Millenium Bridge London
 Pedestrian induced vibration on the opening day
 (synchronization)



ULB

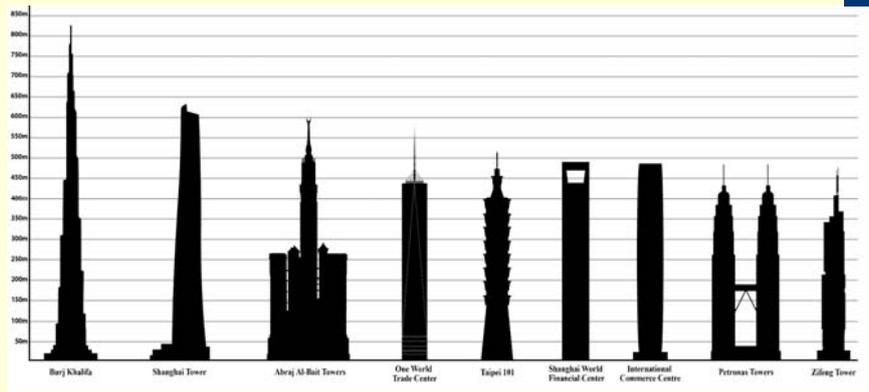
6

Synchronization of metronomes
[Synchronization of pendulum clocks was first observed by Huygens (17^e)]

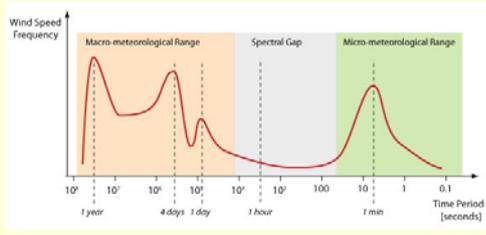


ULB 7

Tallest buildings in the world



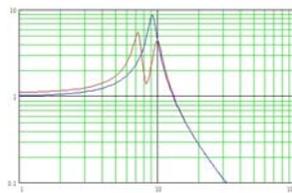
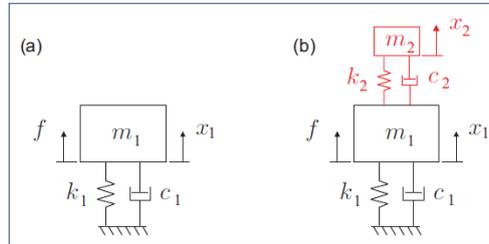
ULB



$T_1 \sim 0.1 n$

8

Reduction of resonance peaks with passive devices:
Tuned Mass Damper (TMD) Dynamic Vibration Absorber (DVA)



Den Hartog optimum design

$$\frac{\omega_2}{\omega_1} = \frac{1}{1 + \mu}$$

$$\xi_2 = \sqrt{\frac{3\mu}{8(1 + \mu)}}$$

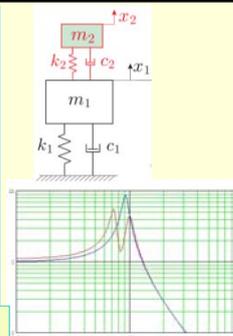
ULB

9

Taipei 101 (509 m)



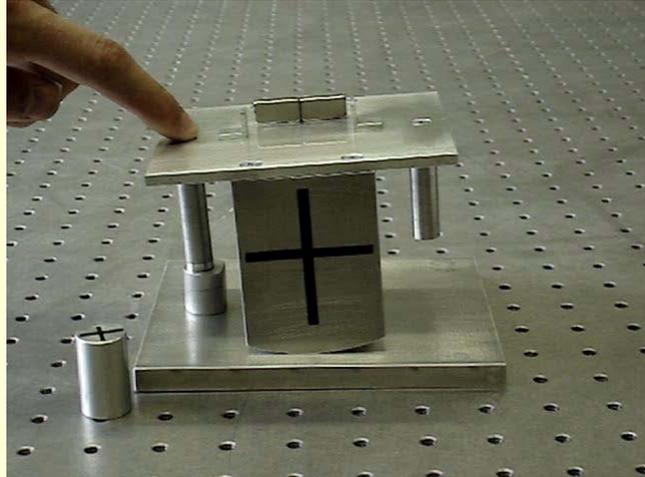
730 T Tuned Mass Damper



ULB

$T_1 \sim 0.1 n$

10

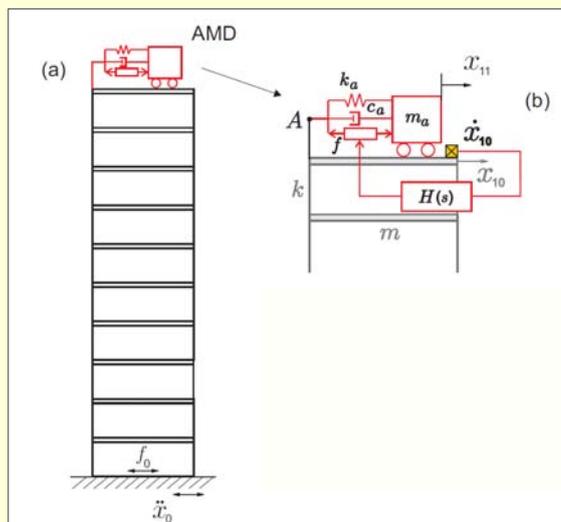


ULB

11

ULB

Active Mass Damper

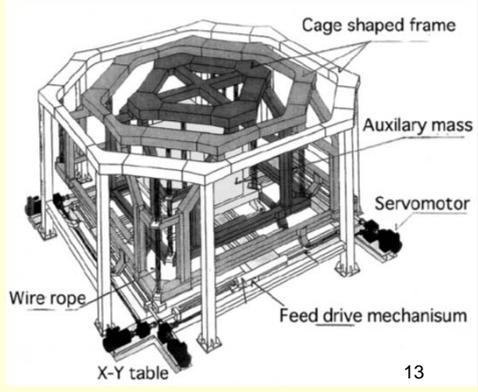
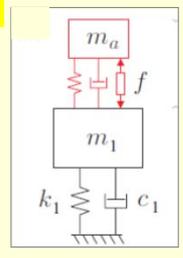


12

Yokohama Landmark Tower
(296 m, 1993)



Active Mass Damper



ULB

13

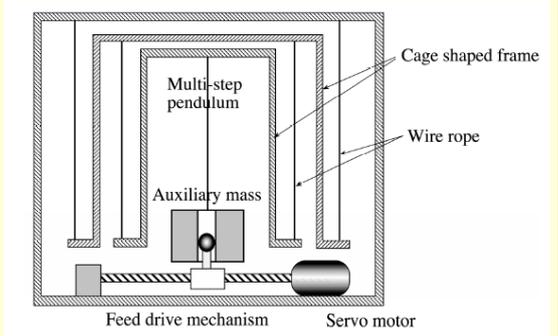
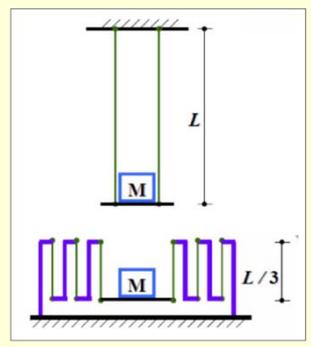


Fig. 2.7 Multi-step pendulum type HMD

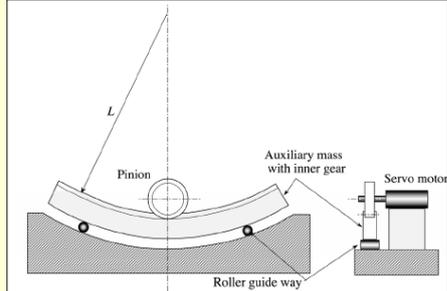
(from K. Seto)

ULB

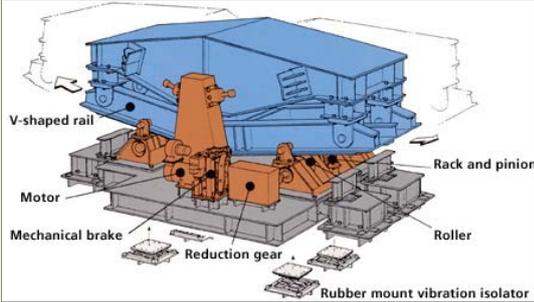
14

ULB

Roller guide HMD



V-shaped HMD

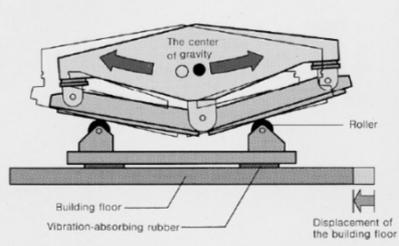


15

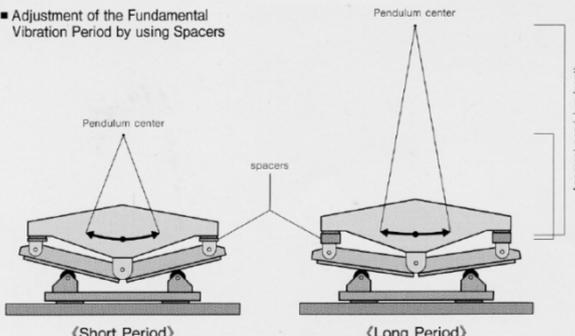
ULB

**TRIGON pendulum device (Kajima corp. 1993)
Shinjuku Park Tower – Seismic Response Control**

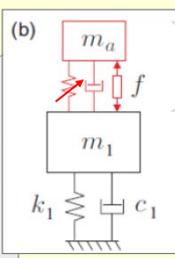
■ The TRIGON Weight Moves Like a Pendulum



■ Adjustment of the Fundamental Vibration Period by using Spacers



(b)



16

Suspension bridges

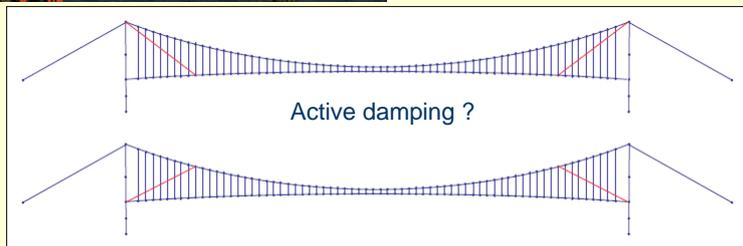
Golden Gate, San Francisco, 1280 m, 1937



Clifton Bridge Bristol (Brunel, 1864)



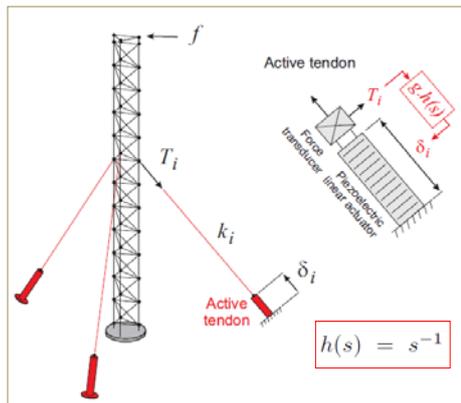
Largest span: Akashi bridge, Kobe 1991 m, 1998



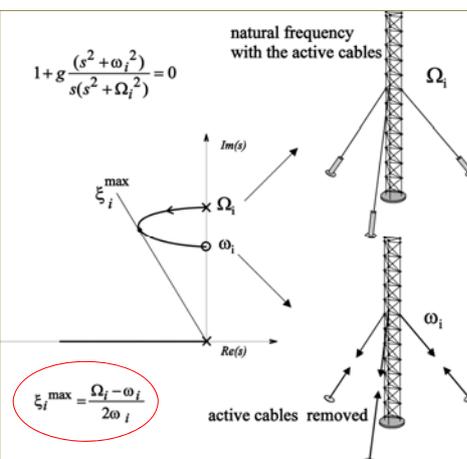
ULB

17

Decentralized active tendon control of cable-structures



Maximum achievable damping:



$$\xi_i^{\max} = \frac{\Omega_i - \omega_i}{2\omega_i}$$

ULB

A.PREUMONT: *Vibration Control of Active Structures*, 3rd Edition, Springer 2011

18

Suspension bridge: Seriate footbridge



ULB

19

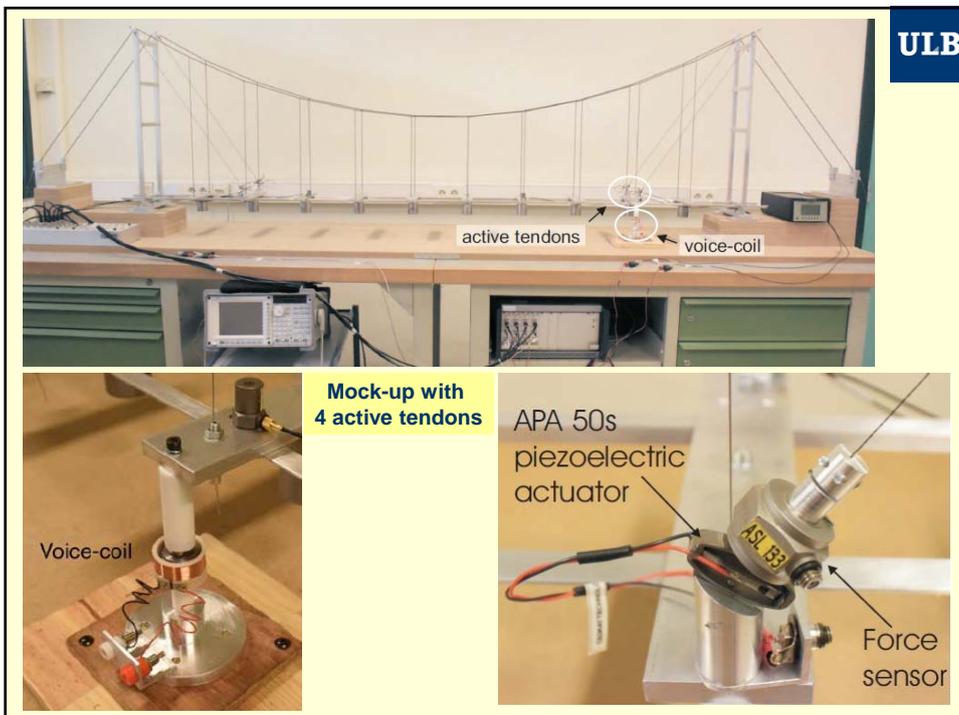
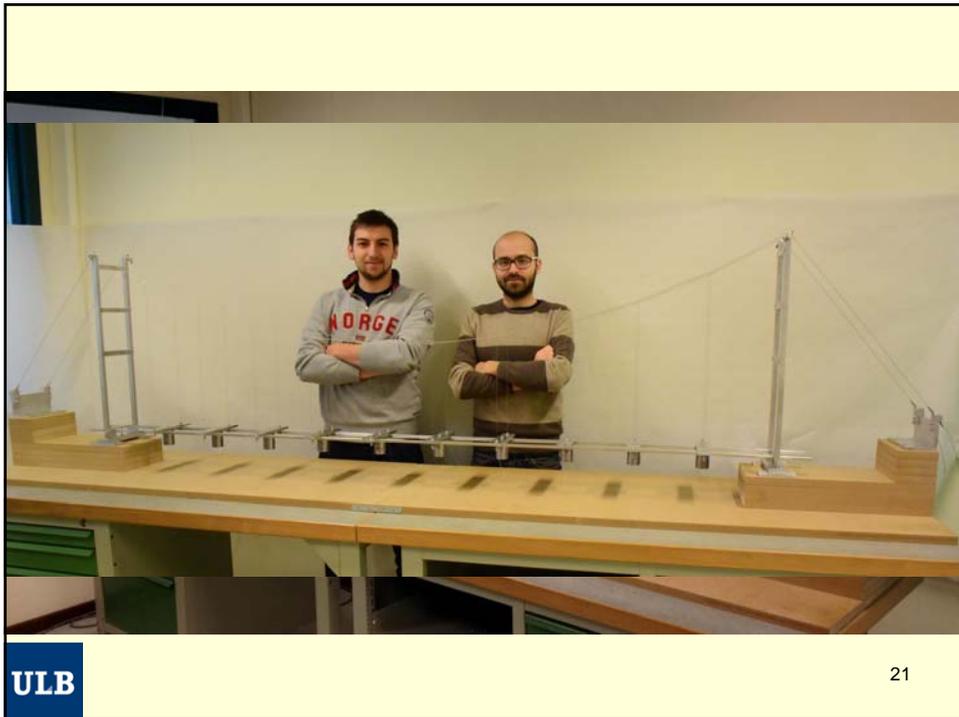
Mode N.	2D Numerical (Hz)	3D Numerical (Hz)	Experimental (Hz)	Numerical Mode Shape	Experimental Mode Shape
1 st B.	1.03	1.02	1.03 $\xi_1 = 2.77\%$		
2 nd B.	1.39	1.48	1.48 $\xi_2 = 1.34\%$		
1 st T.	/	1.79	1.92		
2 nd T.	/	2.1	1.94		
3 rd B.	2.22	2.20	2.17 $\xi_3 = 1.48\%$		
3 rd T.	/	2.65	2.75		
4 th B.	2.81	2.78	2.86 $\xi_4 = 1.50\%$		

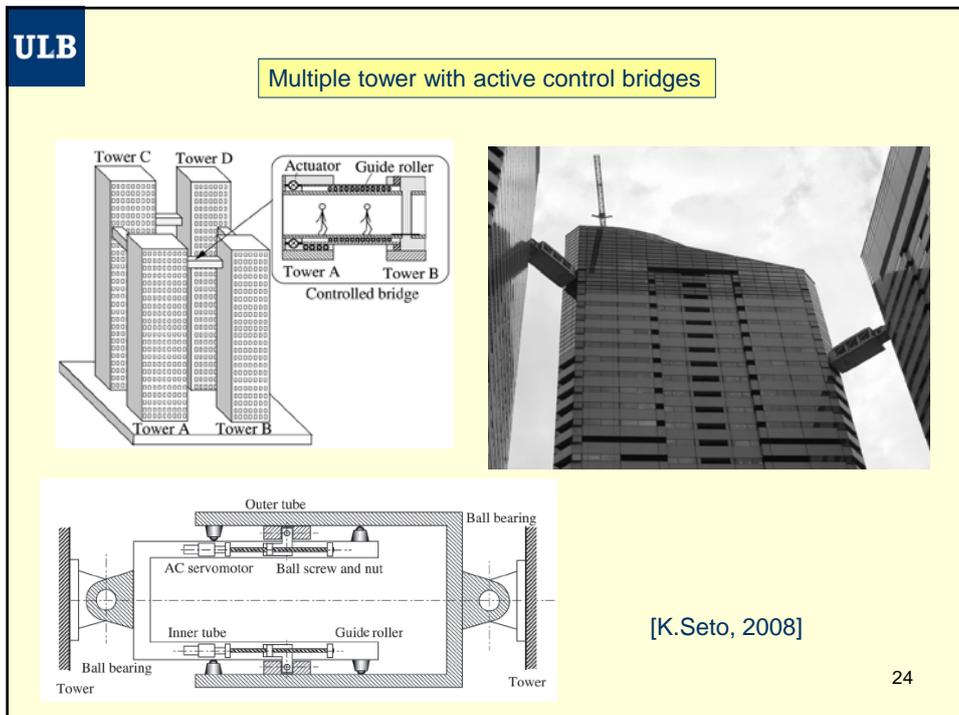
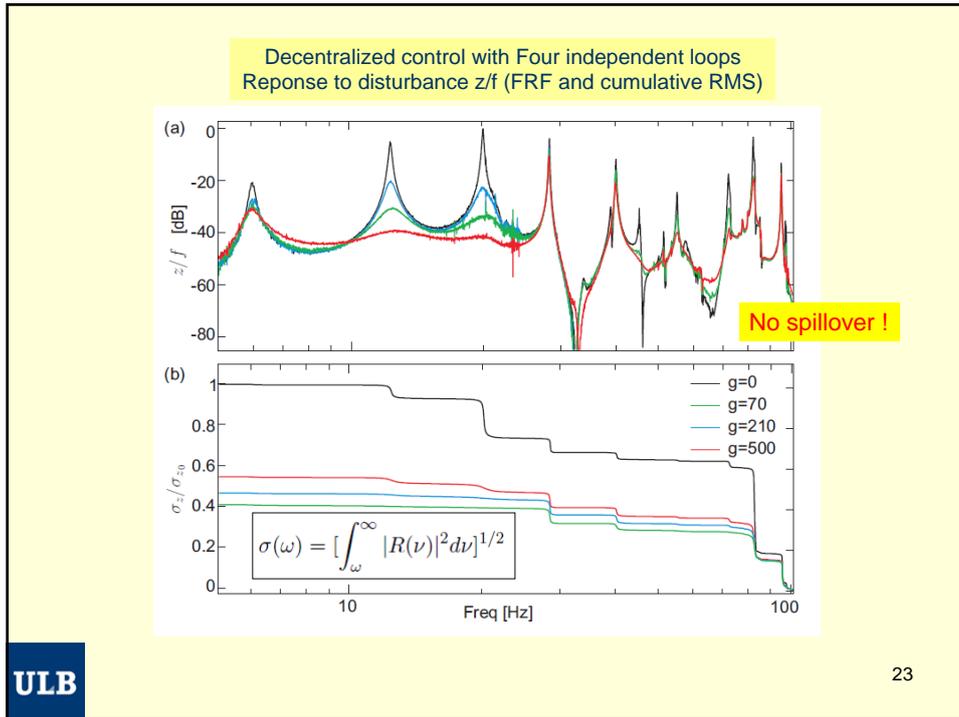
Table 1: Natural frequencies and mode shapes of the Seriate footbridge, comparison of the 3D model and 2D model with experiments [18]. The two critical modes are 3B and 4B.

Active tendon control with 4 steel cables of diameter 10 mm can bring up to 15% damping in the critical modes

ULB

20





Multiple tower with active control
Bridges - Laboratory Mock-up

[K.Seto, 2008]

ULB

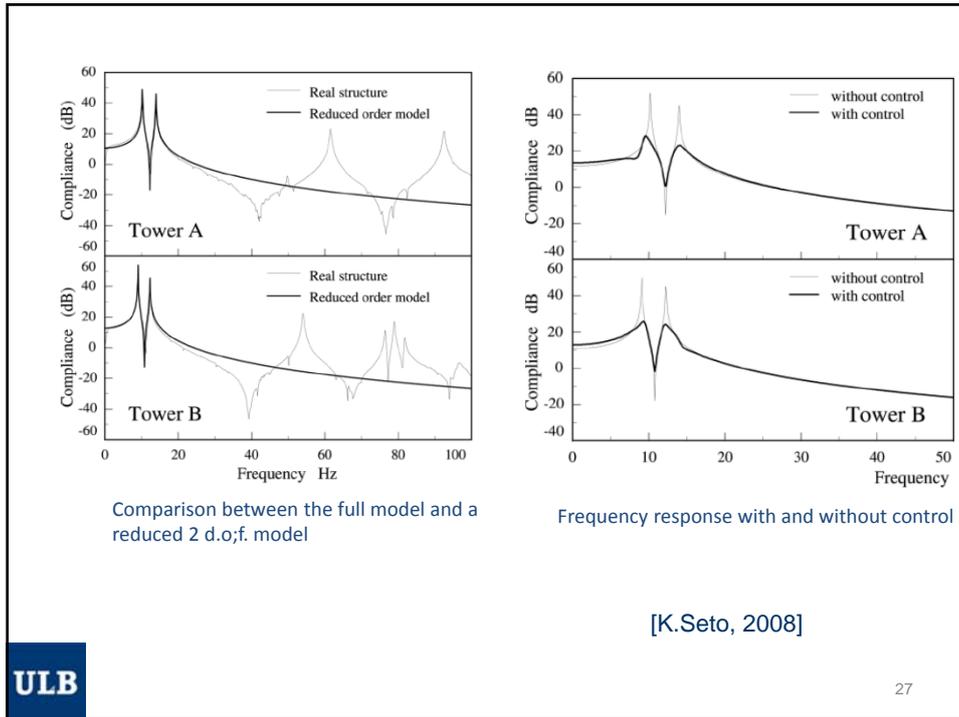
25

Shaking table tests

[K.Seto, 2008]

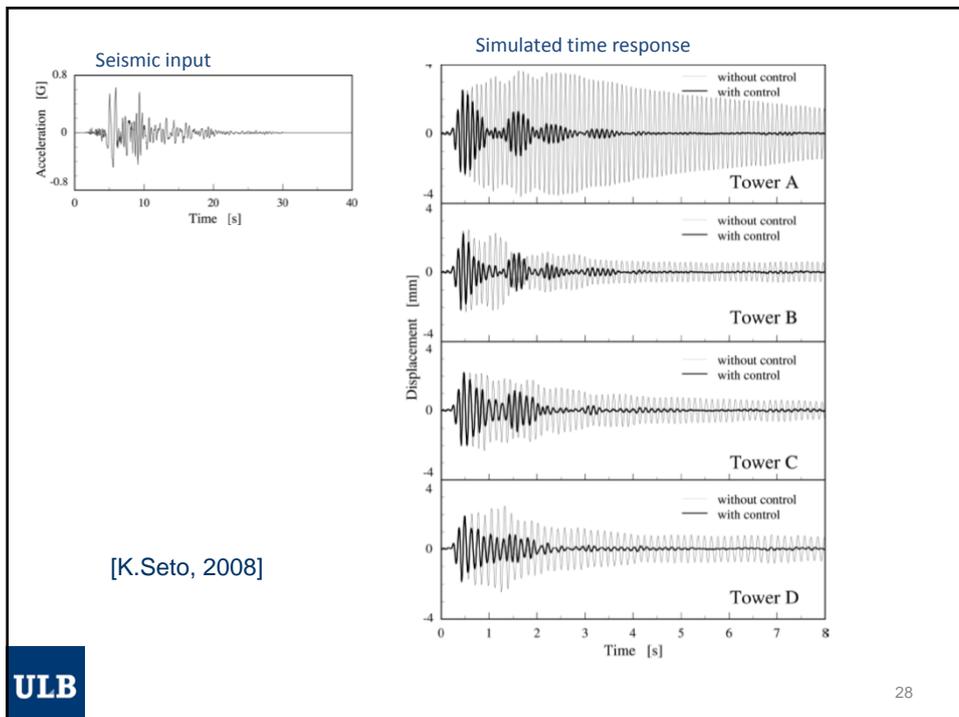
ULB

26



ULB

27

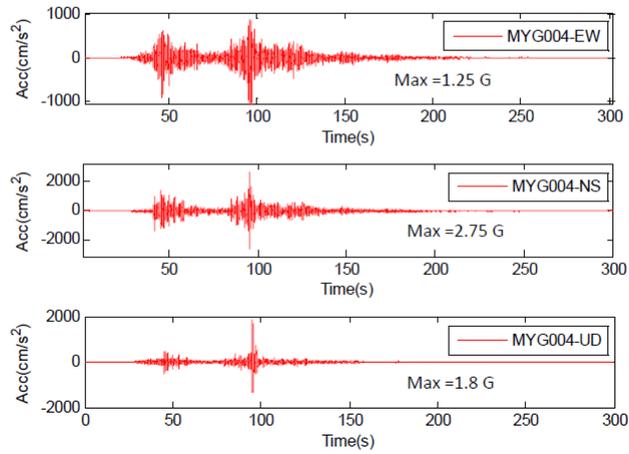


ULB

28

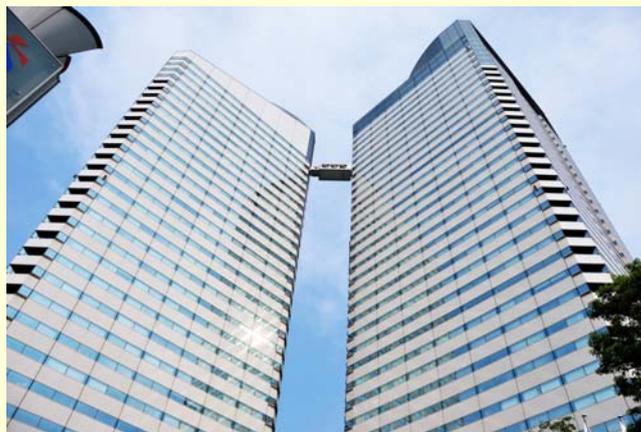
March 11, 2011 Earthquake (Fukushima)

Records with largest PGA : Tsukidate Station, Miyagi. Distance to epicenter : 125.9 Km, Distance to Fault 75Km



[Y.Fujino, 2011]

Triton Square building during Japon earthquake (3/2011)



ULB

Space interferometer

Independent pointing telescopes

Laser metrology

Vibration isolator

Large truss

Beam combiner

Attitude Control

delay line

Vibration isolator

Optical Path Difference (OPD) accuracy: $\lambda/20$

- Vibration isolation from attitude control system & cryocooler
- Damping large truss structures
- Co-phasing telescopes (delay lines)
- Wavefront Control (Adaptive Optics)

31

James Webb Space Telescope (~ 2018)

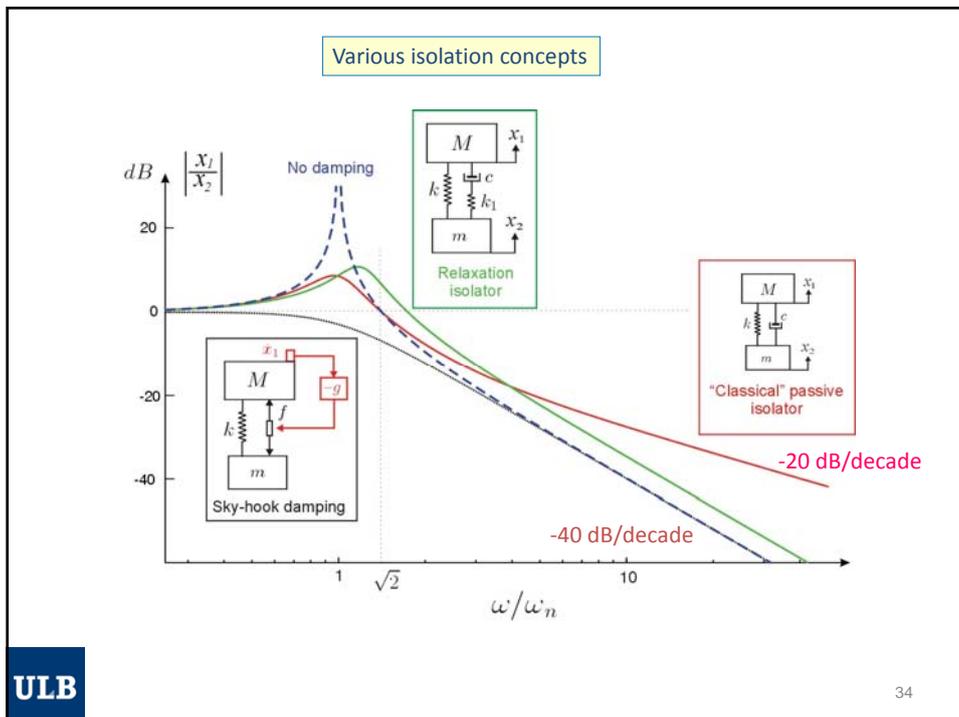
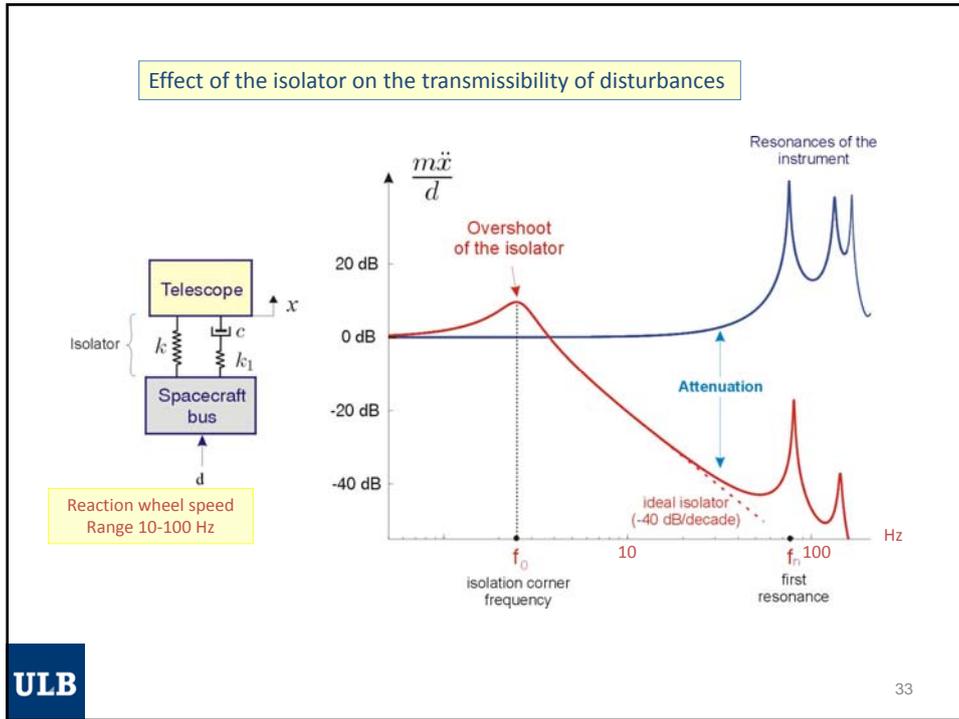
$\varnothing = 6.6m$

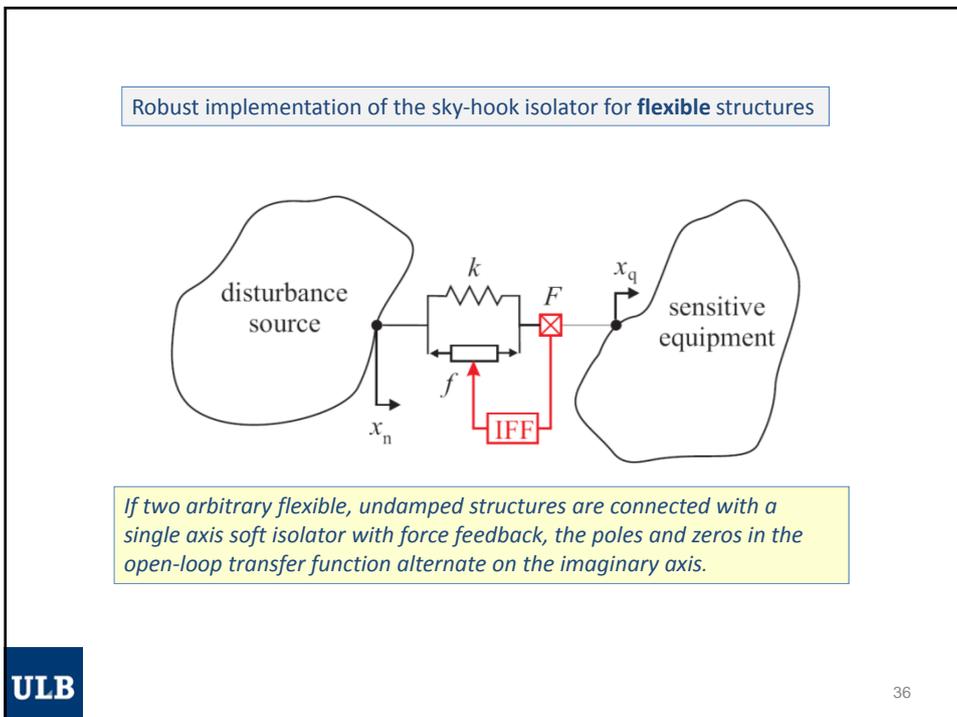
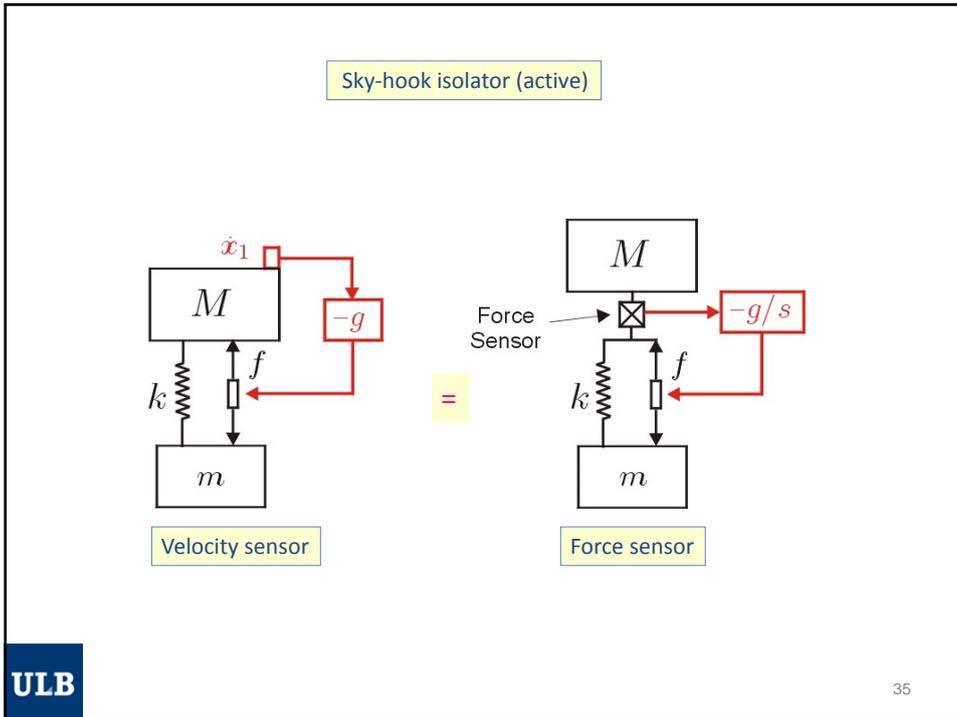
1 Hz Isolator

RWA Isolator 7 Hz

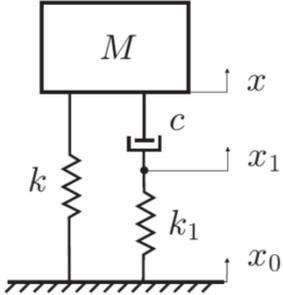
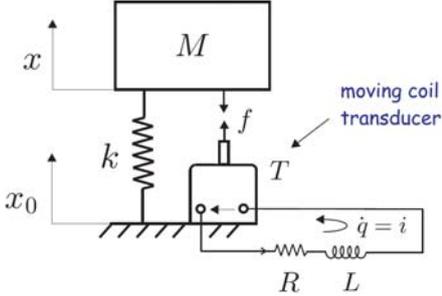
ULB

32





Electromagnetic implementation of the relaxation isolator

The two systems are equivalent provided that:

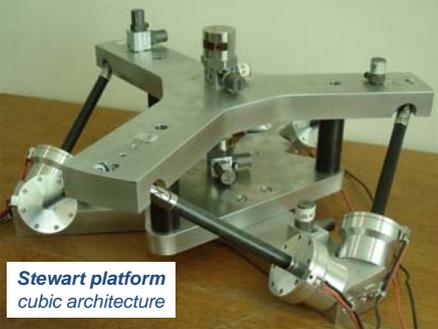
$$k_1 = \frac{T^2}{L} \quad c = \frac{T^2}{R}$$

Electric components are often more flexible and less temperature sensitive than elastomers

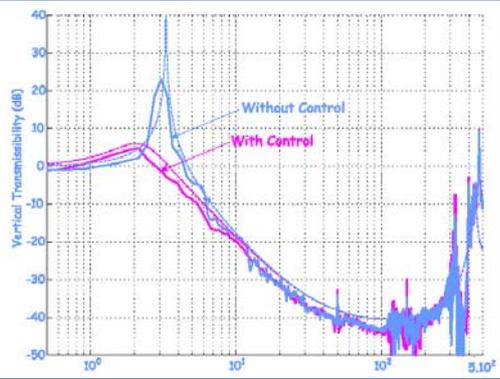

37

Six-axis isolator Zero-g experiment

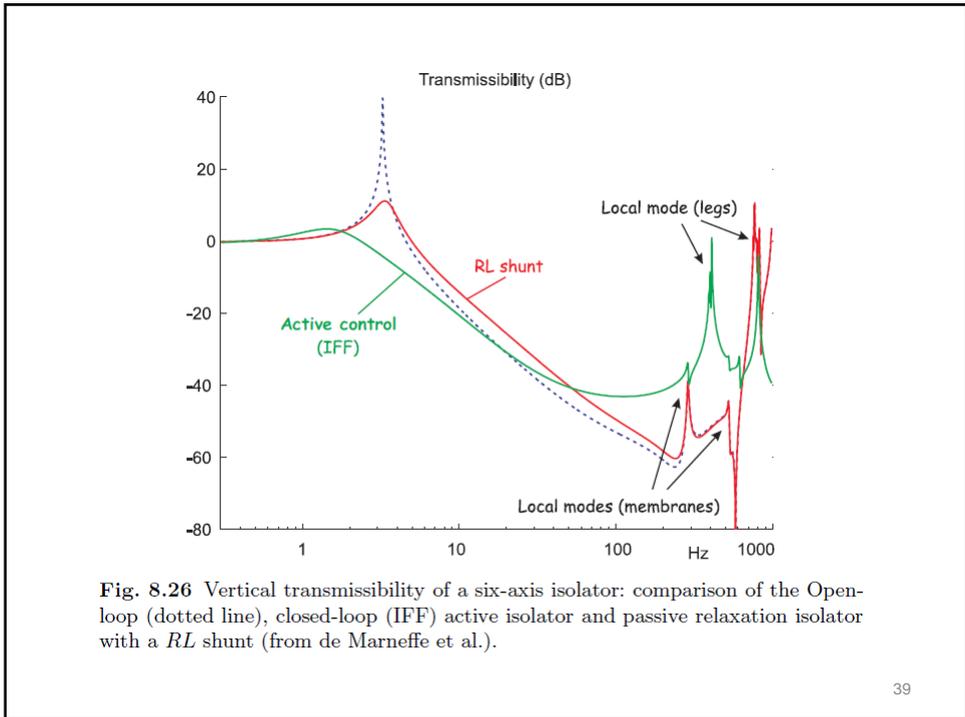




Stewart platform
cubic architecture




38

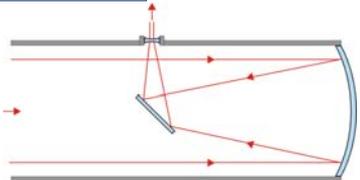


A brief history of earth-based telescopes

Galileo – 1610
discovered Jupiter's Moons
(refractive telescope)

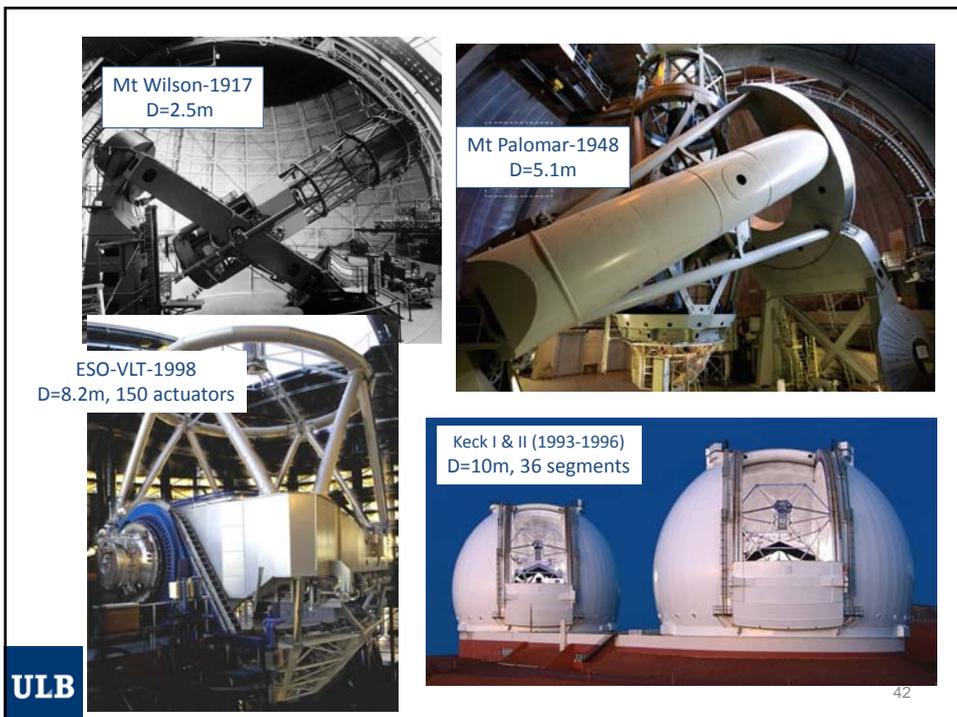
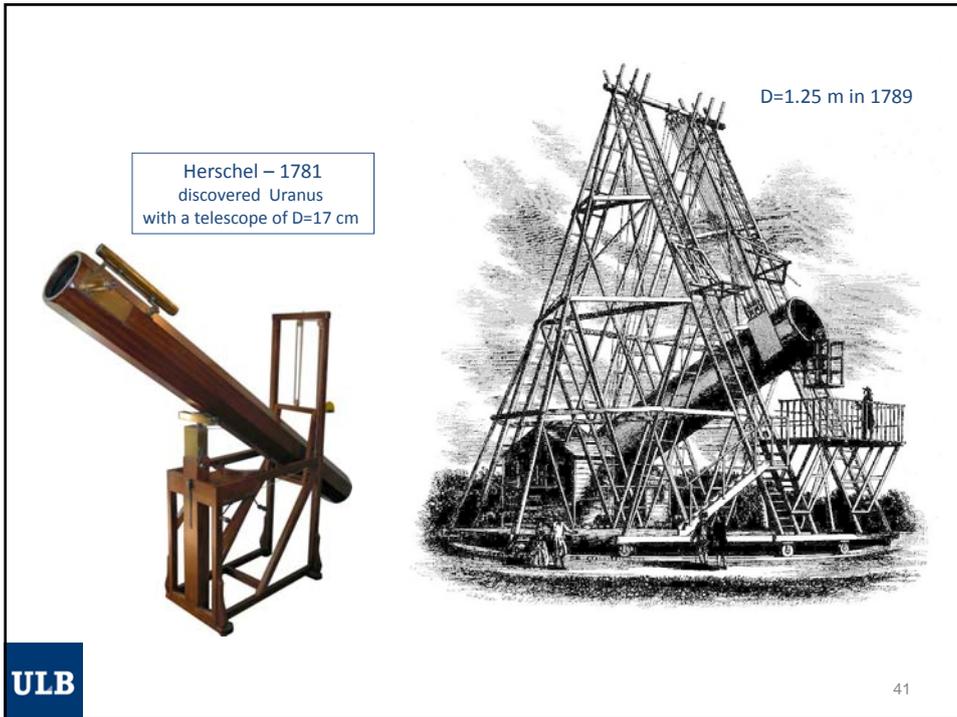


Newton – 1668
reflective telescope
(removes the chromatic aberration)



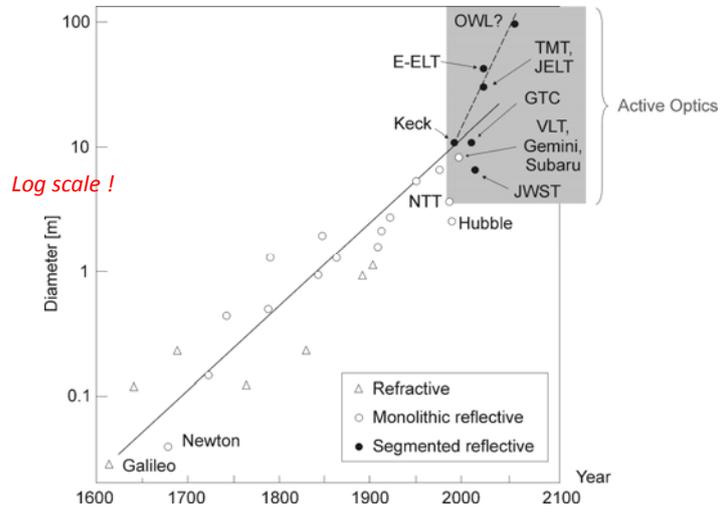
ULB

40



How big can it be ?

Size of optical telescopes vs. time



ULB

43

ULB

M1 of existing and future Optical telescopes

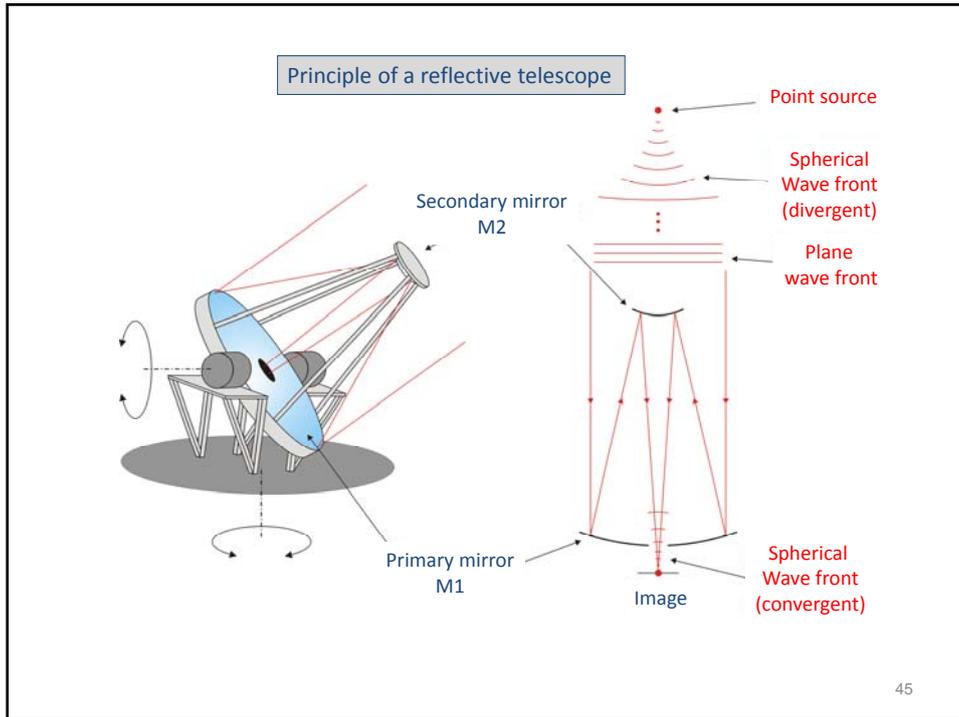
Space telescopes

HST – 2.6m	Herschel – 3.5m	JWST – 6.5m (2018 ?)
------------	-----------------	----------------------

Ground based telescopes

VLT – 8m	Keck – 10m	TMT – 30m (2018 ?)	E-ELT – 39m (2017 ?)
----------	------------	--------------------	----------------------

44



Optical aberrations

Motivation for larger M1: Image quality

$d_1 \sim \lambda/D$

Point source

Diffraction

Aberration

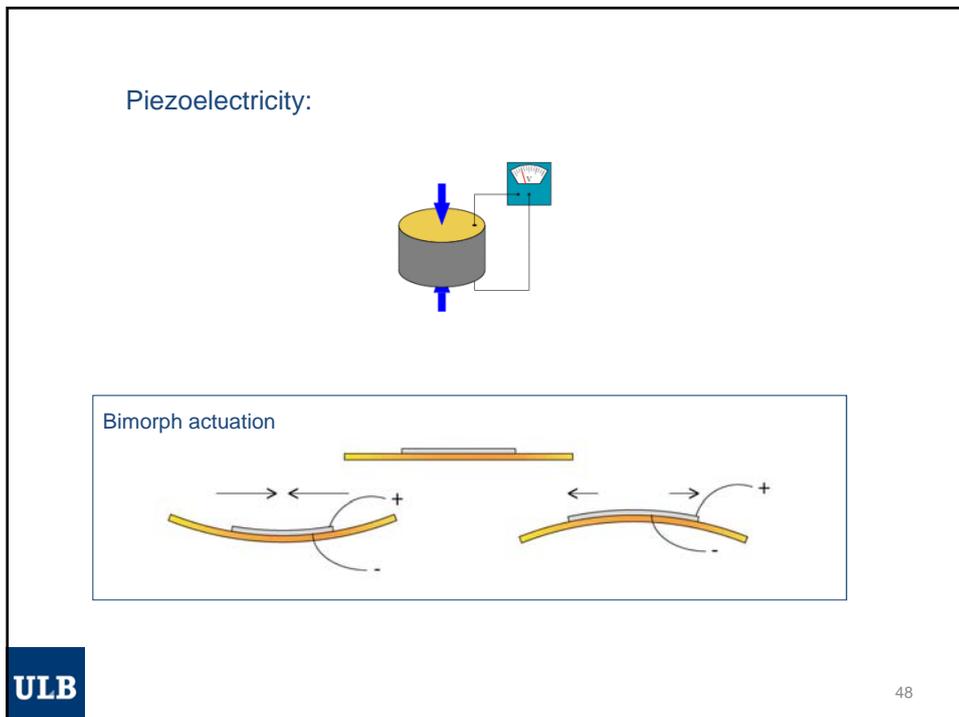
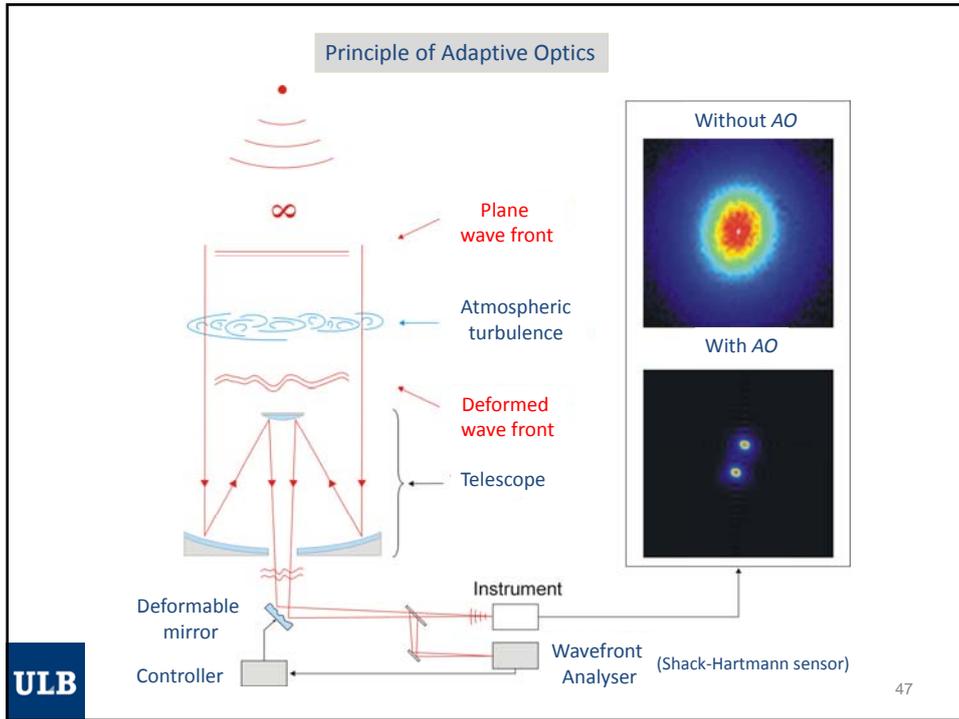
Sources of aberrations (ground based telescopes):

Atmospheric turbulence: **Adaptive optics**

Gravity, wind excitation, thermal gradients: **Active optics**

If the RMS wave front error is $< \lambda/14$, the telescope is considered as diffraction limited

46



Bimorph Adaptive Optics Silicon Mirror

Reflecting side Of the mirror

Silicon wafer (700 μm)

Gold electrodes

PZT Thick film deposition (80 microns)

ULB

Source: G. Rodrigues PhD thesis (2010)
Manufacturing: Fraunhofer IKTS (Dresden)

49

Adaptive Optics Bench

Shack-Hartmann
101 x 101 micro-lenses

Voltage Amplifiers

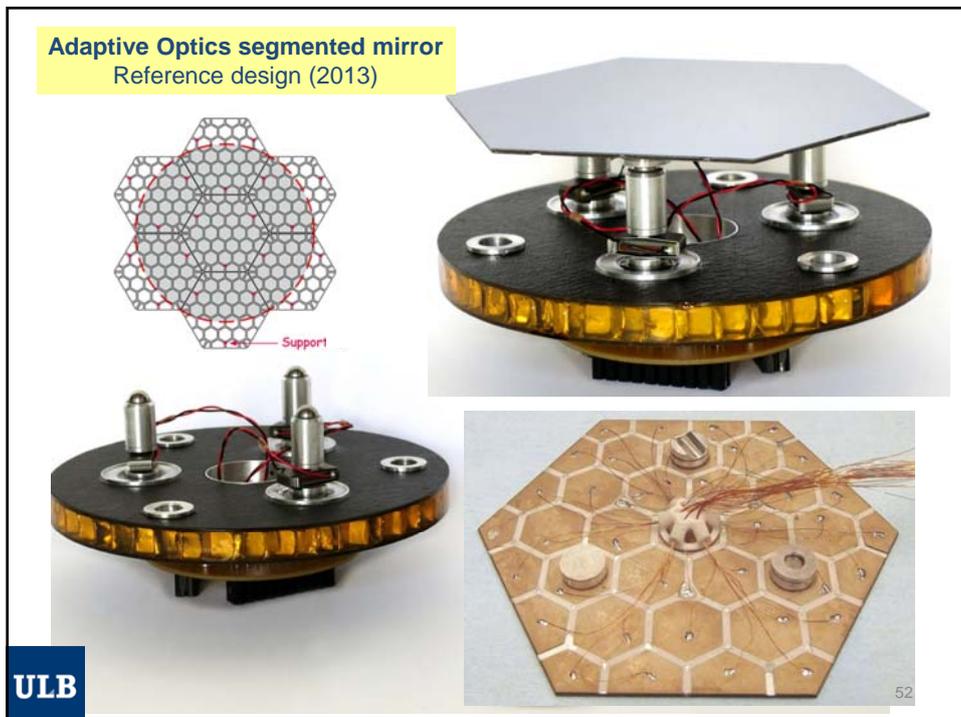
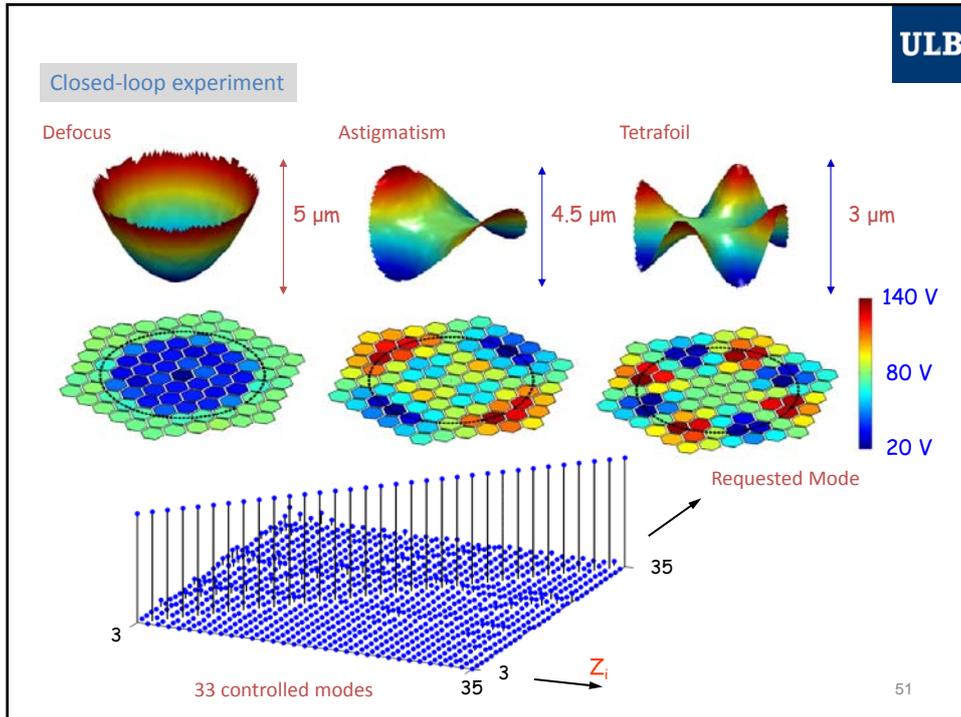
Bimorph Mirror

Telescope

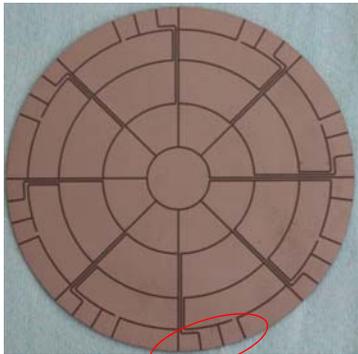
ULB

Collaboration with CSL

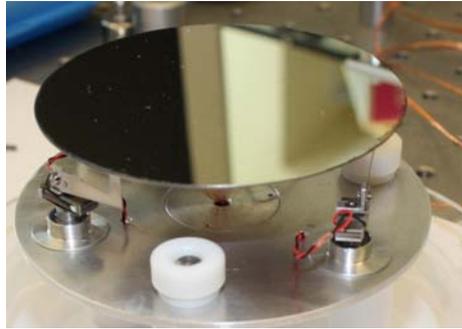
50



AO mirror with PZT actuation
Keystone electrode design (ESA-2015)



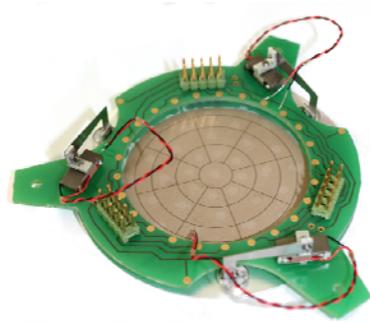
All electrical connections
Outside the pupil



Rigid body actuators

ULB

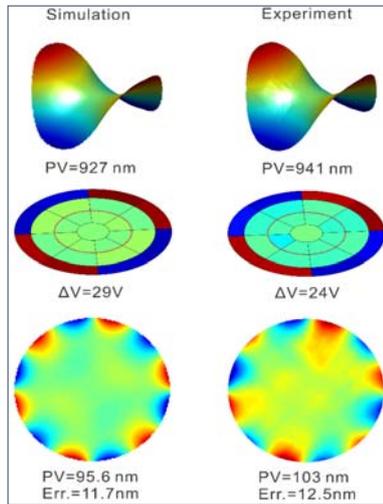
Ultra-flat design (2016)



ULB

54

Comparison of numerical simulation and experiment for an incremental surface of 1 μ m of astigmatism. From top to bottom: surface figure, voltage map and residual error.

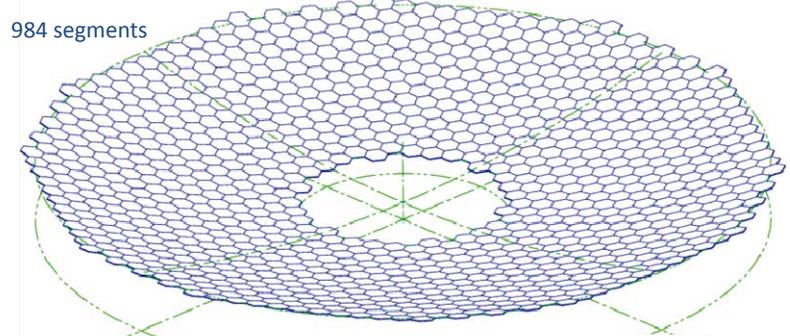


ULB

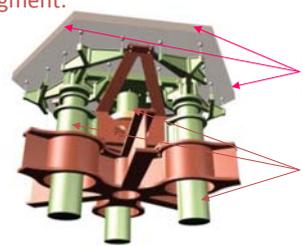
55

Active Optics

E-ELT primary mirror [ESO-2018?]

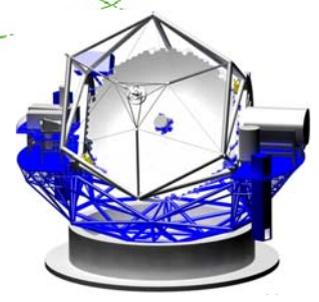


One segment:

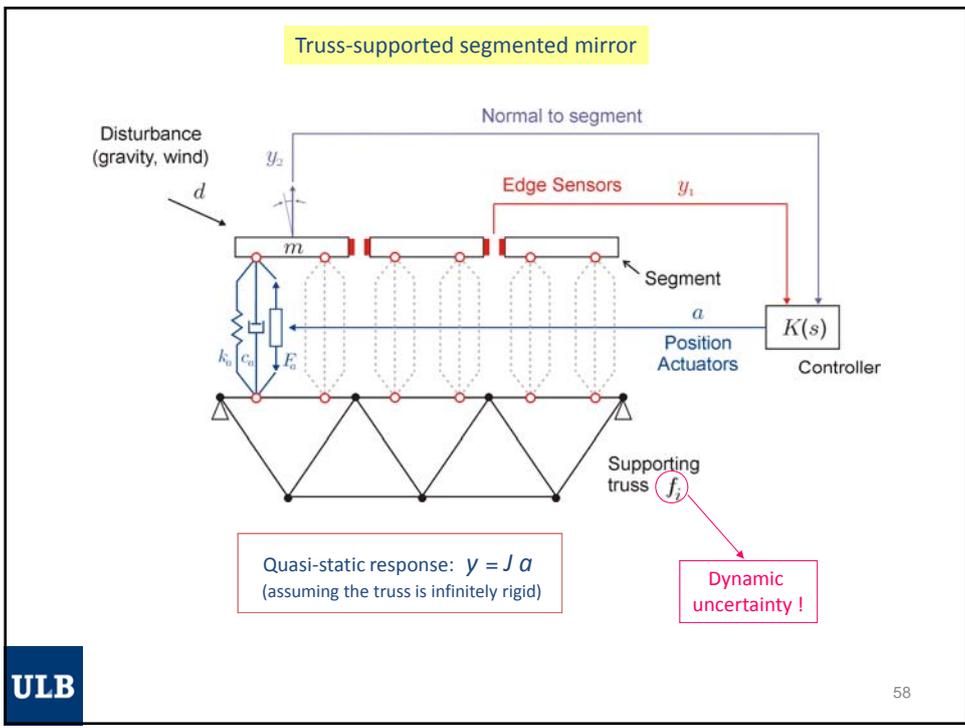
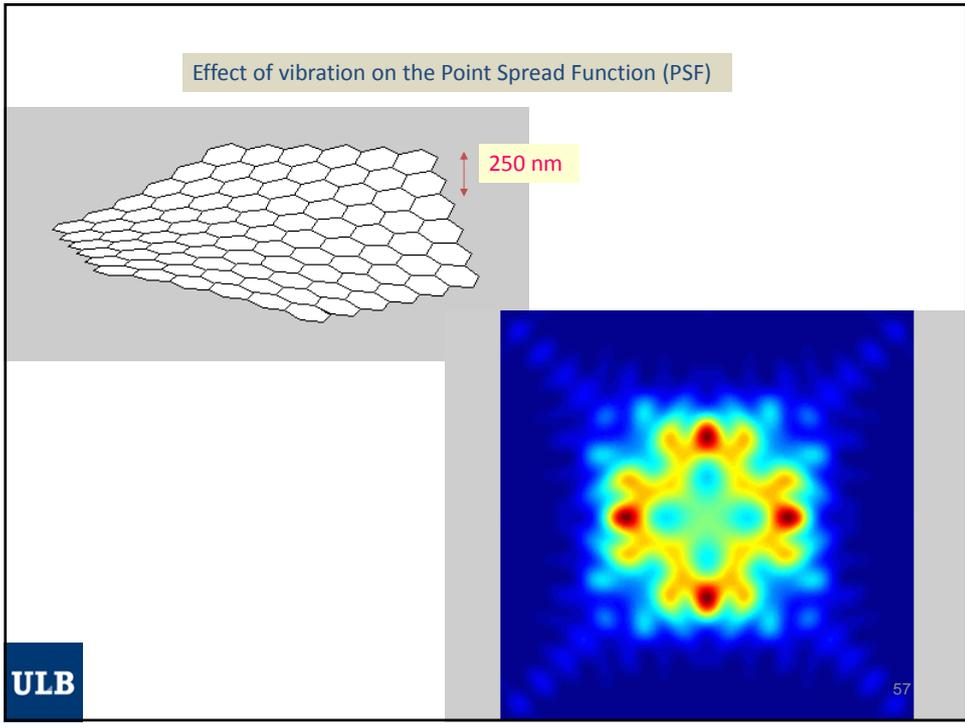


6 Edge sensors

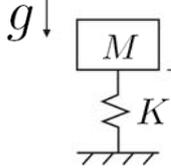
3 Position actuators (two stages)



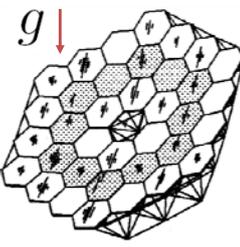
ULB



Static deflection under gravity



$$\Delta = \frac{Mg}{K} = \frac{g}{\omega_1^2}$$



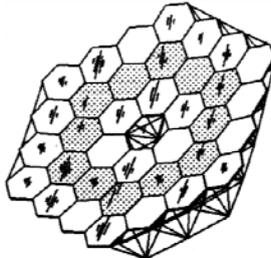
The static deflections under gravity scale according to

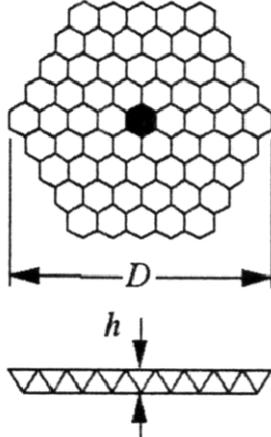
$\Delta \propto f_1^{-2}$

ULB

59

Truss supported segmented reflector
Scaling law for the first natural frequency





$$f_1 \sim \frac{0.852}{D} \left(\frac{h}{D}\right) \sqrt{\eta \cdot \left(\frac{E}{\rho}\right)}$$

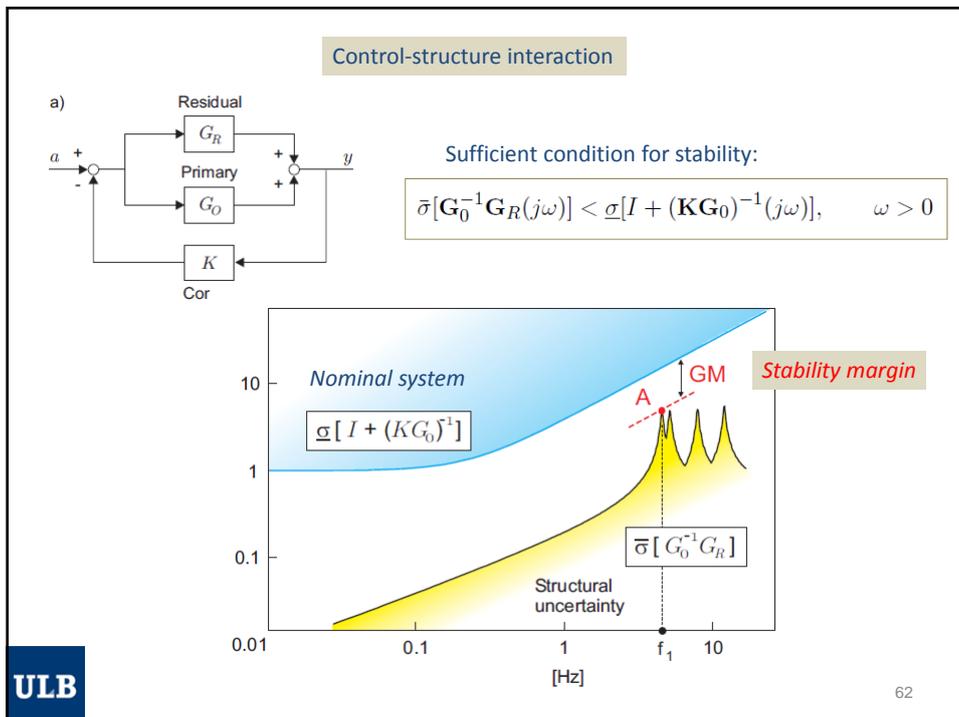
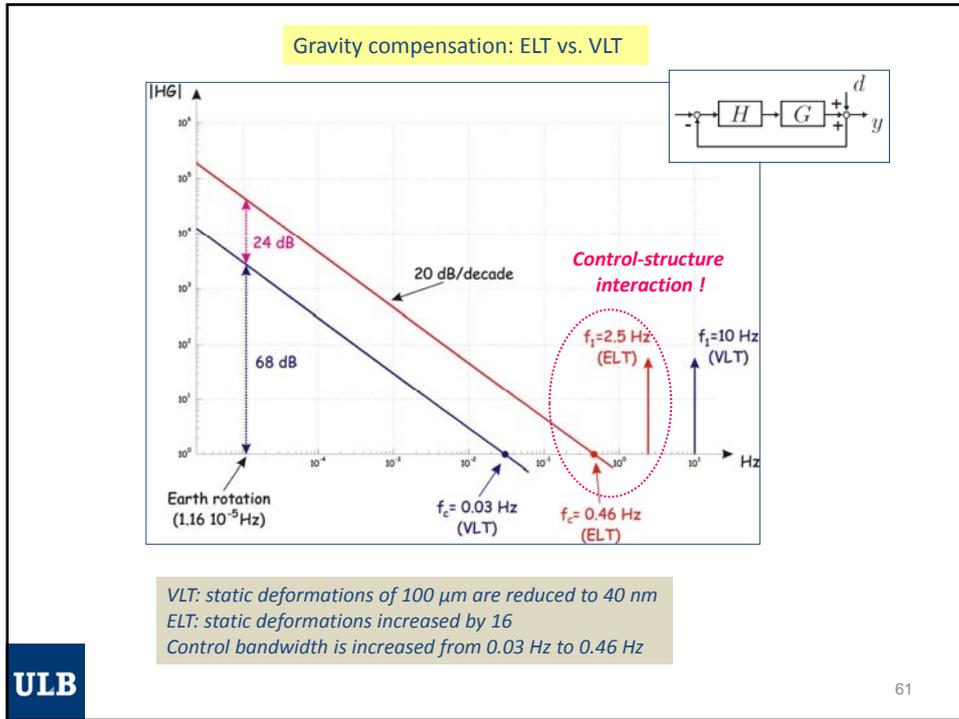
Specific modulus

$$\eta = \frac{\text{Truss Mass}}{\text{Truss Mass} + \text{Reflector Mass}}$$

ULB

(Source: Lake, Peterson & Mikulas, 2006)

60



Conclusions:

- The **size** of civil engineering structures increases rapidly
- They become **more sensitive to vibrations** (wind, earthquake, traffic)
- This constitutes new opportunities for **active vibration control**
- The **actuator** is the most important component of the control system

Acknowledgements:

All my PhD students:

<http://www.ulb.ac.be/scmero/publi.html>

Our sponsors:

EU, ESA, ESF, FNRS-FRIA, Région Wallonne, FCT Portugal, Humboldt foundation...

Many colleagues scattered all over Europe and overseas,
from whom I have learned my job...